

# Introduction to Mathematical Micromagnetics

Valeriy Slastikov



## Lecture 2. Basic structures: 2D boundary vortex

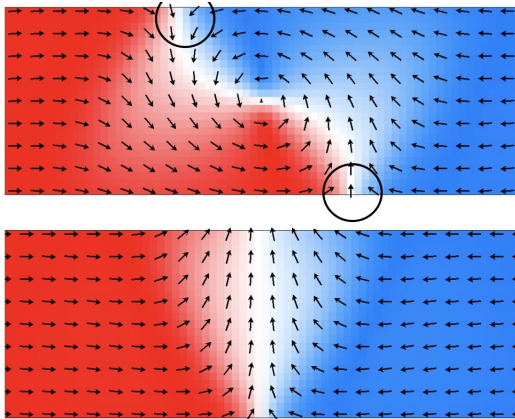


Figure: Vortex and transverse DWs (Jamet et al., (2015))

# Basic structures: 2D boundary vortex

We want to understand a structure of a *boundary vortex*. The simplified energy is ( $\mathbf{m} \in \mathbb{S}^1$ )

$$E(\mathbf{m}) = \varepsilon^2 \int_{\mathbb{R}_+^2} |\nabla \mathbf{m}(\mathbf{r})|^2 d^2 r + \int_{\mathbb{R}} m_2^2(0, x) dx. \quad (1.1)$$

We cannot talk about minimizers in a usual sense as for vortex boundary conditions

$$\mathbf{m}(\mathbf{r}) \rightarrow \frac{(-y, x)}{|\mathbf{r}|} \text{ as } |\mathbf{r}| \rightarrow \infty. \quad (1.2)$$

energy is infinite. Instead we want to look at critical points.

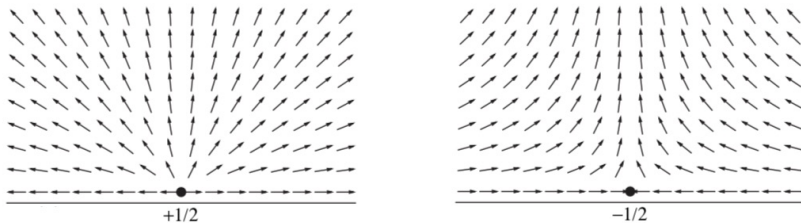


Figure: Boundary vortices (Tchernyshyov, Chern, *PRL*, (2005))

## 2D boundary vortex - reformulation

We can reformulate the problem as a scalar one and link it to 1D non-local gradient phase transitions. Indeed, using  $\mathbf{m} = (\cos \theta, \sin \theta)$  the energy becomes

$$E(\mathbf{m}) \equiv F(\theta) = \varepsilon^2 \int_{\mathbb{R}_+^2} |\nabla \theta|^2 d^2 r + \int_{\mathbb{R}} \sin^2 \theta d\mathcal{H}^1. \quad (1.3)$$

The critical points/local minimizers of this energy have been extensively studied (e.g. Toland, *J. Func. An.*, (1997); Cabre, Sola-Morales, *CPAM*, (2005); Kurzke, *CVPDE*, (2006); Moser, *ARMA*, (2004) )

Let us link the energy (1.3) with some nonlocal 1D energy. We can fix a trace of  $\theta$  at the boundary to be a given function,  $\theta(x, 0) = \bar{\theta}(x)$  and then reformulate the problem in terms of  $\bar{\theta}$ . It amounts to obtain a harmonic extension of  $\bar{\theta}$  on a half-plane, or solve

$$\Delta \theta = 0 \text{ in } \mathbb{R}_+^2, \quad \theta(x, 0) = \bar{\theta}(x). \quad (1.4)$$

We can then plug this solution into the energy and minimize with respect to  $\bar{\theta}$ .

## 2D boundary vortex - nonlocal 1D energy

Let us proceed informally and assume  $\bar{\theta} \in C_c^\infty(\mathbb{R})$ . We can define the Fourier transform in  $x$  variable as

$$\hat{\theta}(k, y) = \int_{\mathbb{R}} e^{-ikx} \theta(x, y) dx. \quad (1.5)$$

We also note that  $\hat{\theta}(k, 0) = \hat{\bar{\theta}}(k)$ . Writing the Euler-Lagrange equations (recall that  $\bar{\theta}$  is fixed) we obtain

$$-\partial_y^2 \hat{\theta}(k, y) + k^2 \hat{\theta}(k, y) = 0, \quad \hat{\theta}(k, 0) = \hat{\bar{\theta}}(k), \quad \partial_y \hat{\theta}(k, \infty) = 0. \quad (1.6)$$

The solution of this equation is  $\hat{\theta}(k, y) = \hat{\bar{\theta}}(k) e^{-ky}$ .

We can now rewrite our energy using Fourier representation as

$$F(\theta) := \varepsilon^2 \int_0^\infty \int_{\mathbb{R}} |\partial_y \hat{\theta}(k, y)|^2 + k^2 |\hat{\theta}(k, y)|^2 dk dy + \int_{\mathbb{R}} \sin^2 \bar{\theta}(x) dx \quad (1.7)$$

## 2D boundary vortex - nonlocal 1D energy

We plug expression for  $\hat{\theta}(k, y)$  into the energy and obtain

$$F(\bar{\theta}) = \varepsilon^2 \int_{\mathbb{R}} \int_0^{\infty} 2k^2 e^{-2ky} |\hat{\theta}(k)|^2 dy dk + \int_{\mathbb{R}} \sin^2 \bar{\theta}(x) dx \quad (1.8)$$

$$= \int_{\mathbb{R}} k |\hat{\theta}(k)|^2 dk + \int_{\mathbb{R}} \sin^2 \bar{\theta}(x) dx. \quad (1.9)$$

Taking inverse Fourier transform we obtain

$$F(\bar{\theta}) = \frac{\varepsilon^2}{\pi} \int_{\mathbb{R}} \int_{\mathbb{R}} \frac{|\bar{\theta}(x) - \bar{\theta}(y)|^2}{|x - y|^2} dx dy + \int_{\mathbb{R}} \sin^2 \bar{\theta}(x) dx. \quad (1.10)$$

We observe that our energy is now informally represented as a non-local gradient phase transition problem

$$F(\bar{\theta}) = \varepsilon^2 \int_{\mathbb{R}} |\partial^{\frac{1}{2}} \bar{\theta}(x)|^2 dx + \int_{\mathbb{R}} \sin^2 \bar{\theta}(x) dx. \quad (1.11)$$

Various aspects of similar problems have been extensively studied (e.g.

Alberti, Bouchitté, Seppecher, *C.R. Acad. Sci. Paris*, (1994); Palatucci, Savin, Valdinoci, *Ann. Mat. Pure Appl.*, (2013) )

## 2D boundary vortex - classification of critical points

The Euler-Lagrange equation is

$$\frac{\varepsilon^2}{\pi} \int_{\mathbb{R}} (2\bar{\theta}(x) - \bar{\theta}(x - \xi) - \bar{\theta}(x + \xi)) \frac{1}{|\xi|^2} d\xi + \sin 2\bar{\theta}(x) = 0 \quad \forall x \in \mathbb{R}.$$

We can explicitly check that

$$\bar{\theta}(x) = \frac{\pi}{2} \pm \arctan(x/\varepsilon^2) + \pi n \quad (1.12)$$

is a non-trivial solution of this equation.

The Euler-Lagrange equation for original problem is

$$\Delta\theta = 0 \text{ in } \mathbb{R}_+^2, \quad \varepsilon^2 \partial_y \theta(x, 0) = -\frac{1}{2} \sin 2\theta(x, 0). \quad (1.13)$$

We can classify solutions of this problem (Toland, *J. Func. An.*, (1997))

### Theorem 1

Let  $\theta$  be a bounded solution of (1.13). Then either  $\theta$  is constant solution, periodic solution, or there exists  $a \in \mathbb{R}$ ,  $n \in \mathbb{Z}$  such that

$$\theta(x, y) = \frac{\pi}{2} \pm \arctan \frac{x+a}{y+\varepsilon^2} + \pi n. \quad (1.14)$$

## Basic structures: 2D baby skyrmion

We want to understand a skyrmion structure in thin films. The simplified energy is

$$E(\mathbf{m}) = \int_{\mathbb{R}^2} |\nabla \mathbf{m}(\mathbf{r})|^2 d^2 r \quad (1.15)$$

defined in the class

$$\mathcal{A} = \{\mathbf{m} \in H_{loc}^1(\mathbb{R}^2; \mathbb{R}^3) : \nabla \mathbf{m} \in L^2(\mathbb{R}^2; \mathbb{R}^3 \times \mathbb{R}^3), |\mathbf{m}(x)| = 1 \text{ a.e.}\}. \quad (1.16)$$

We impose *skyrmion boundary conditions*

$$\mathbf{m}(\mathbf{r}) \rightarrow -\mathbf{e}_3 \text{ as } |\mathbf{r}| \rightarrow \infty \quad (1.17)$$

and a topological degree condition

$$\mathcal{N}(\mathbf{m}) = \frac{1}{4\pi} \int_{\mathbb{R}^2} (\partial_1 \mathbf{m} \times \partial_2 \mathbf{m}) \cdot \mathbf{m} d^2 r = 1. \quad (1.18)$$



## Basic structures: 2D baby skyrmion

It is clear that

$$E(\mathbf{m}) = \int_{\mathbb{R}^2} |\nabla \mathbf{m}(\mathbf{r})|^2 d^2r = \int_{\mathbb{R}^2} |\partial_2 \mathbf{m} - \mathbf{m} \times \partial_1 \mathbf{m}|^2 d^2r + 8\pi \geq 8\pi \quad (1.19)$$

We want to find a suitable profile such that

$$\partial_2 \mathbf{m} - \mathbf{m} \times \partial_1 \mathbf{m} = 0 \quad (1.20)$$

This is Belavin-Polyakov profile (Belavin, Polyakov, *JETP Lett.* (1975))

$$\mathbf{m} = \left( \frac{\pm 2x}{1+r^2}, \frac{\pm 2y}{1+r^2}, \frac{1-r^2}{1+r^2} \right). \quad (1.21)$$

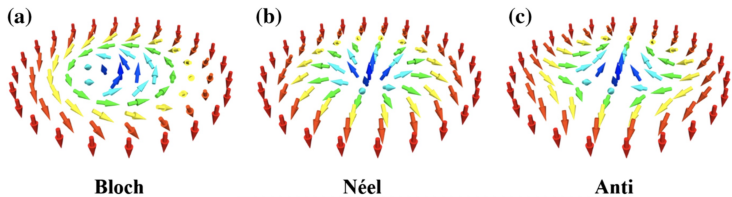


Figure: (a) Bloch skyrmion; (b) Neel skyrmion; (c) Anti-skyrmion

# Basic structures: 2D baby skyrmion.

We can formulate a general problem. Consider the energy

$$E(\mathbf{m}) = \int_{\Omega} |\nabla \mathbf{m}(\mathbf{r})|^2 d^2r - 2\kappa \int_{\Omega} \nabla m_3 \cdot \mathbf{m} d^2r \\ + (Q - 1) \int_{\Omega} (1 - m_3^2) d^2r + 2h \int_{\Omega} (1 + m_3) d^2r,$$

where  $\mathbf{m} \in \mathcal{A}$ ,  $\Omega \subseteq \mathbb{R}^2$ ,  $\mathbf{m} = -\mathbf{e}_3$  on  $\mathbb{R}^2 \setminus \Omega$ ,  $\mathcal{N}(\mathbf{m}) = d \in \mathbb{Z}$ .

- Existence of minimizers for  $d = 1$  (Melcher, *Proc. R. Soc. Lon. A* (2014); Monteil, Muratov, Simone, S, *Com. Math. Phys.* (2023))
- Existence of minimizers for  $d \neq \{0, 1\}$ ? **partial results**
- What are minimizing profiles? (Li, Melcher, *J. Func. An.* (2018))
- Existence/structure of skyrmion lattice? (Hill, S, Tchernyshev, *SciPost Phys.* (2021))

## Lecture 2. Magnetostatic energy.

The magnetostatic / stray field energy is defined as

$$E_{ms}(\mathbf{M}) = -\frac{\mu_0}{2} \int_{\Omega} \mathbf{H}_d \cdot \mathbf{M} d^3r. \quad (1.22)$$

Here  $\mathbf{H}_d$  is a demagnetizing field solving

$$\nabla \cdot (\mathbf{H}_d + \mathbf{M}) = 0, \quad \nabla \times \mathbf{H}_d = 0 \quad (1.23)$$

Introduce a scalar potential  $U_d : \mathbb{R}^3 \rightarrow \mathbb{R}$  with  $\mathbf{H}_d = -\nabla U_d$  and  $U_d$  solves

$$\int_{\mathbb{R}^3} \nabla U_d \cdot \nabla \phi d^3r = \int_{\Omega} \mathbf{M} \cdot \nabla \phi d^3r, \quad \text{for any } \phi \in C_c^\infty(\mathbb{R}^3). \quad (1.24)$$

The stray field energy can be rewritten as

$$E_{ms}(\mathbf{M}) = \frac{\mu_0}{2} \int_{\Omega} \nabla U_d \cdot \mathbf{M} d^3r = \frac{\mu_0}{2} \int_{\mathbb{R}^3} |\nabla U_d|^2 d^3r. \quad (1.25)$$

Solving Poisson equation we obtain

$$E_{ms}(\mathbf{M}) = \frac{\mu_0}{8\pi} \int_{\mathbb{R}^3} \int_{\mathbb{R}^3} \frac{\nabla \cdot \mathbf{M}(\mathbf{r}) \nabla \cdot \mathbf{M}(\mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|} d^3r d^3r'. \quad (1.26)$$

# Magnetostatic energy: Min/Max problems.

Rescaling  $\mathbf{M}$  to  $\mathbf{m}$ ,  $U_d$  to  $u_m$  and  $E_{ms}$  we can represent

$$E_{ms}(\mathbf{m}) = \frac{1}{2} \int_{\mathbb{R}^3} |\nabla u_m|^2 d^3 r, \quad \Delta u_m = \nabla \cdot \mathbf{m}. \quad (1.27)$$

## Theorem 2

For any  $\mathbf{m} \in L^2(\Omega; \mathbb{R}^3)$  with  $\Omega \subset \mathbb{R}^3$  bounded there is unique solution  $u_m \in H^1(\mathbb{R}^3)$  of the following maximization problem

$$\max_{u \in H^1(\mathbb{R}^3)} \int_{\Omega} \mathbf{m} \cdot \nabla u d^3 r - \frac{1}{2} \int_{\mathbb{R}^3} |\nabla u|^2 d^3 r. \quad (1.28)$$

Moreover,  $u_m$  satisfies Euler-Lagrange equations

$$\int_{\mathbb{R}^3} \nabla u_m \cdot \nabla \phi d^3 r = \int_{\Omega} \mathbf{m} \cdot \nabla \phi d^3 r, \quad \text{for any } \phi \in C_c^\infty(\mathbb{R}^3) \quad (1.29)$$

and

$$E_{ms}(\mathbf{m}) = \frac{1}{2} \int_{\mathbb{R}^3} |\nabla u_m|^2 d^3 r = \max_{u \in H^1(\mathbb{R}^3)} \int_{\Omega} \mathbf{m} \cdot \nabla u d^3 r - \frac{1}{2} \int_{\mathbb{R}^3} |\nabla u|^2 d^3 r.$$

## Magnetostatic energy: vector potential.

There is a way to represent magnetostatic energy through minimization problem. We have Maxwell's equations

$$\nabla \cdot (\mathbf{h}_d + \mathbf{m}) = 0, \quad \nabla \times \mathbf{h} = 0. \quad (1.30)$$

Before we used  $\mathbf{h}_d = -\nabla u$  to obtain maximization problem using potential  $u$ . Now we can use  $\mathbf{h}_d + \mathbf{m} = \nabla \times \mathbf{a}$  with Coulomb gauge  $\nabla \cdot \mathbf{a} = 0$ . This leads to

$$\nabla \times (\nabla \times \mathbf{a}_m) = -\Delta \mathbf{a}_m = \nabla \times \mathbf{m}. \quad (1.31)$$

The magnetostatic energy is

$$E_{ms} = \frac{1}{2} \int_{\mathbb{R}^3} |\mathbf{h}_d|^2 d^3 r = \frac{1}{2} \int_{\mathbb{R}^3} |\nabla \times \mathbf{a}_m - \mathbf{m}|^2 d^3 r. \quad (1.32)$$

# Magnetostatic energy: Min/Min problem.

We can formulate a minimization problem (Di Fratta et. al, *SIMA*, (2020))

## Theorem 3

For any  $\mathbf{m} \in L^2(\Omega; \mathbb{R}^3)$  with  $\Omega \subset \mathbb{R}^3$  bounded there is unique solution  $\mathbf{a}_m \in \dot{H}^1(\mathbb{R}^3)$  of the following minimization problem

$$\min_{\mathbf{a} \in \dot{H}^1(\mathbb{R}^3; \mathbb{R}^3)} \frac{1}{2} \int_{\mathbb{R}^3} |\nabla \mathbf{a}|^2 + \frac{1}{2} \int_{\Omega} |\mathbf{m}|^2 - \int_{\Omega} \mathbf{m} \cdot \nabla \times \mathbf{a}. \quad (1.33)$$

Moreover,  $\mathbf{a}_m$  satisfies Euler-Lagrange equations

$$-\Delta \mathbf{a}_m = \nabla \times \mathbf{m} \quad \text{in } \dot{H}^{-1}(\mathbb{R}^3; \mathbb{R}^3) \quad (1.34)$$

and

$$E_{ms}(\mathbf{m}) = \min_{\mathbf{a} \in \dot{H}^1(\mathbb{R}^3; \mathbb{R}^3)} \frac{1}{2} \int_{\mathbb{R}^3} |\nabla \mathbf{a}|^2 + \frac{1}{2} \int_{\Omega} |\mathbf{m}|^2 - \int_{\Omega} \mathbf{m} \cdot \nabla \times \mathbf{a}.$$

Useful for localized upper bounds and as a double minimization

$$\min_{\mathbf{m}} \min_{\mathbf{a}} E(\mathbf{m}, \mathbf{a}) = \min_{\mathbf{a}} \min_{\mathbf{m}} E(\mathbf{m}, \mathbf{a}) \quad (1.35)$$

# Magnetostatic energy - two representations

1. We define  $n$  - dimensional Fourier transforms as

$$\mathcal{F}(f)(\mathbf{k}) \equiv \hat{f}(\mathbf{k}) = \int_{\mathbb{R}^n} f(\mathbf{r}) e^{-i\mathbf{r}\cdot\mathbf{k}} d^n\mathbf{r}. \quad (1.36)$$

Using equation for magnetostatic potential  $u$  we obtain  $\hat{u}(\mathbf{k}) = \frac{\hat{\mathbf{m}}(\mathbf{k})\cdot\mathbf{k}}{|\mathbf{k}|^2}$

$$\int_{\mathbb{R}^3} |\nabla u|^2 d^3\mathbf{r} = \frac{1}{(2\pi)^3} \int_{\mathbb{R}^3} |\widehat{\nabla u}|^2 d^3\mathbf{k} = \frac{1}{(2\pi)^3} \int_{\mathbb{R}^3} \frac{|\mathbf{k}\cdot\hat{\mathbf{m}}(\mathbf{k})|^2}{|\mathbf{k}|^2} d^3\mathbf{k}. \quad (1.37)$$

2. Energy can be written using  $u(\mathbf{r}) = -\frac{1}{4\pi} \int_{\Omega} \frac{\nabla\cdot\mathbf{m}(\mathbf{r}')}{|\mathbf{r}-\mathbf{r}'|} + \frac{1}{4\pi} \int_{\partial\Omega} \frac{(\mathbf{n}\cdot\mathbf{m})(\mathbf{r}')}{|\mathbf{r}-\mathbf{r}'|}$  as

$$E_{ms}(\mathbf{m}) \equiv \frac{1}{2} \int_{\Omega} \nabla u \cdot \mathbf{m} = \int_{\Omega} \int_{\Omega} \frac{\nabla\cdot\mathbf{m}(\mathbf{r}) \nabla\cdot\mathbf{m}(\mathbf{r}')}{8\pi|\mathbf{r}-\mathbf{r}'|} \quad (1.38)$$
$$+ \int_{\partial\Omega} \int_{\partial\Omega} \frac{(\mathbf{n}\cdot\mathbf{m})(\mathbf{r}) (\mathbf{n}\cdot\mathbf{m})(\mathbf{r}')}{8\pi|\mathbf{r}-\mathbf{r}'|} - 2 \int_{\Omega} \int_{\partial\Omega} \frac{\nabla\cdot\mathbf{m}(\mathbf{r}) (\mathbf{n}\cdot\mathbf{m})(\mathbf{r}')}{8\pi|\mathbf{r}-\mathbf{r}'|}$$

# Ferromagnetic thin films

We consider the following thin film domain

$$\Omega_\varepsilon = \{(x, y, z) : z \in [0, \varepsilon], (x, y) \in \omega \subset \mathbb{R}^2\}, \quad (1.39)$$

where  $0 < \varepsilon \ll 1$  and want to simplify magnetostatic energy.

**Fact.** Let  $\tilde{\mathbf{m}} = \frac{1}{\varepsilon} \int_0^\varepsilon \mathbf{m}(x, y, z) dz$  and let  $\Delta \bar{u} = \nabla \cdot \tilde{\mathbf{m}}$ . Then the following inequality holds:

$$\left| \int_{\mathbb{R}^3} |\nabla u|^2 - \int_{\mathbb{R}^3} |\nabla \bar{u}|^2 \right| \leq C \varepsilon^{\frac{3}{2}} \|\partial_z \mathbf{m}\|_{L^2(\Omega_\varepsilon)}. \quad (1.40)$$

This allows us to remove dependence of  $\mathbf{m}$  on  $z$  variable in magnetostatics.

We can now use Fourier representation and after calculation obtain

$$\begin{aligned} \int_{\mathbb{R}^3} |\nabla \bar{u}|^2 &= \frac{\varepsilon}{(2\pi)^2} \int_{\mathbb{R}^2} |\widehat{m}_3(\mathbf{k}')|^2 d^2 \mathbf{k}' + \frac{\varepsilon}{(2\pi)^2} \int_{\mathbb{R}^2} |\widehat{\mathbf{m}'}(\mathbf{k}') \cdot \mathbf{k}'|^2 \frac{1 - \hat{\Gamma}_\varepsilon(|\mathbf{k}'|)}{|\mathbf{k}'|^2} d^2 \mathbf{k}' \\ &\quad - \frac{\varepsilon}{(2\pi)^2} \int_{\mathbb{R}^2} |\widehat{m}_3(\mathbf{k}') \mathbf{k}'|^2 \frac{1 - \hat{\Gamma}_\varepsilon(|\mathbf{k}'|)}{|\mathbf{k}'|^2} d^2 \mathbf{k}', \end{aligned} \quad (1.41)$$

where

$$\hat{\Gamma}_\varepsilon(|\mathbf{k}'|) = \frac{1 - e^{-\varepsilon|\mathbf{k}'|}}{\varepsilon|\mathbf{k}'|} \sim 1 - \frac{\varepsilon|\mathbf{k}'|}{2} \quad (1.42)$$



# Ferromagnetic thin films: Gioia-James regime

We have the following energy

$$E(\mathbf{m}) = \alpha \int_{\Omega_\varepsilon} |\nabla \mathbf{m}(\mathbf{r})|^2 d^2 r + \int_{\mathbb{R}^3} |\nabla u|^2 \quad (1.43)$$

and want to find its  $\Gamma$ -limit as  $\varepsilon \rightarrow 0$ .

It is clear using DCT that if  $m_{3,\varepsilon} \rightarrow m_3$  in  $L^2(\omega; \mathbb{R}^3)$  we have

$$\int_{\mathbb{R}^2} |\widehat{m_{3,\varepsilon}}(\mathbf{k}') \mathbf{k}'|^2 \frac{1 - \widehat{\Gamma}_\varepsilon(|\mathbf{k}'|)}{|\mathbf{k}'|^2} d^2 \mathbf{k}' \rightarrow 0. \quad (1.44)$$

After rescaling domain in  $z$ -variable to

$\Omega = \{(x, y, z) : z \in [0, 1], (x, y) \in \omega \subset \mathbb{R}^2\}$  we obtain

$$E(\mathbf{m}) \sim \varepsilon \alpha \int_{\Omega} |\nabla' \mathbf{m}|^2 + \frac{1}{\varepsilon^2} |\partial_z \mathbf{m}|^2 + \varepsilon \int_{\omega} m_3^2 + o(\varepsilon) \quad (1.45)$$

Therefore, we can show (Gioia, James, *PRSA*, (1997))

$$\frac{1}{\varepsilon} E(\mathbf{m}_\varepsilon) \rightarrow \alpha \int_{\omega} |\nabla \mathbf{m}|^2 + \int_{\omega} m_3^2 \quad (1.46)$$

# Ferromagnetic thin films: Kohn-S regime

We have the following energy

$$E(\mathbf{m}) = \alpha \varepsilon |\ln \varepsilon| \int_{\Omega_\varepsilon} |\nabla \mathbf{m}(\mathbf{r})|^2 d^2 r + \int_{\mathbb{R}^3} |\nabla u|^2 \quad (1.47)$$

and want to find its  $\Gamma$ -limit as  $\varepsilon \rightarrow 0$ . Rescaling domain in  $z$ -variable we obtain

$$\frac{1}{\varepsilon^2 |\ln \varepsilon|} E(\mathbf{m}_\varepsilon) = \alpha \int_{\Omega} |\nabla' \mathbf{m}_\varepsilon|^2 + \frac{1}{\varepsilon^2} |\partial_z \mathbf{m}_\varepsilon|^2 + \frac{1}{\varepsilon^2 |\ln \varepsilon|} \int_{\mathbb{R}^3} |\nabla u|^2 \quad (1.48)$$

We can show that

$$\frac{1}{\varepsilon^2 |\ln \varepsilon|} \int_{\mathbb{R}^3} |\nabla u|^2 \sim \frac{1}{\varepsilon |\ln \varepsilon|} \int_{\omega} m_3^2 + \frac{1}{2\pi} \int_{\partial\omega} (\mathbf{m}_\varepsilon \cdot \mathbf{n})^2 \quad (1.49)$$

And hence (Kohn, S, ARMA, (2005))

$$\frac{1}{\varepsilon^2 |\ln \varepsilon|} E(\mathbf{m}_\varepsilon) \rightarrow \alpha \int_{\omega} |\nabla \mathbf{m}|^2 + \frac{1}{2\pi} \int_{\partial\omega} (\mathbf{m} \cdot \mathbf{n})^2, \quad (1.50)$$

where  $\mathbf{m} = \mathbf{m}(x, y)$  and  $|\mathbf{m}| = 1$ .